

# Quantized Cosmological Constant in (1+1)-dimensional Dilaton-Maxwell gravity

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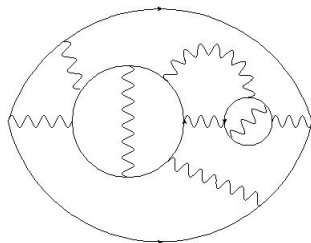
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# The cosmological constant problem

Strict experimental bound:  $|\lambda| \sim 10^{-123}$  in Planck units

Can we predict the value of  $\lambda$ ?

Vacuum energy density in QFT  $\leftrightarrow$  cosmological constant in GR



$$\begin{aligned} \rho_{vac} &\sim \lambda_{\text{cutoff}}^4 \\ &\Downarrow \\ \lambda_{EW} &\sim 100 \text{ GeV} \\ &\Downarrow \\ \rho_{vac}/\lambda &\sim 10^{54} \end{aligned}$$

What about gravity?

# Gravity in lower dimensions

## 1d gravity with matter

- Quantization condition on  $\lambda$
- 1 to 1 correspondence between values for  $\lambda$  and quantum states
- $\lambda$  takes values in the energy spectrum of the matter sector

## (1+1)-dimensional Dilaton Gravity

- Simple (non trivial) model of quantum gravity
- Limit of spherically symmetric GR
- Effective limit of superstring models

# The classical model

(1+1)-dimensional Dilaton Gravity with a non-minimal Maxwell field:

$$S_{DM} = \frac{1}{\kappa} \int_{\mathcal{M}} dx^2 \sqrt{-g} \left( XR - U(X) X_{,\mu} X^{,\mu} - 2V(X) - \frac{1}{4} G(X) F_{\mu\nu} F^{\mu\nu} \right)$$

$\mathcal{M}$  compact with no boundaries

Parametrization for the metric tensor:

$$dx^2 = e^\varphi \left( -\lambda_0 \lambda_1 dt^2 + (\lambda_0 - \lambda_1) dt ds + ds^2 \right)$$

Dilaton dependent conformal transformation:

$$g_{\mu\nu} \rightarrow e^{\chi(X)} g_{\mu\nu}$$

# Equivalence with Liouville Field Theory

Isolating the cosmological constant:  $V(X) = \lambda + \lambda v(X)$

$$\text{Decoupling conditions} \quad \begin{cases} U(X) &= \frac{1}{2\xi^2} - \partial_X \ln(v(X)) \\ G(X) &= \frac{2\xi^2 \Lambda_G e^{X/\xi^2}}{\lambda(1+v(X))} \end{cases}$$

$$\text{New fields} \quad \begin{cases} \bar{Z} &= \xi [\varphi + \chi(X) + \ln(1 + v(X))] + \xi \ln(\lambda) \\ \bar{Y} &= \xi [\varphi + \chi(X) + \ln(1 + v(X)) - X/\xi^2] + \xi \ln(\lambda) \end{cases}$$

Extracting the CC from the fields

$$\bar{Z} = Z + \xi \ln(\lambda) \quad \bar{Y} = Y + \xi \ln(\lambda)$$

# The effective action with the cosmological constant

Effective action

$$S_{\text{eff}} = \int_{\mathcal{M}} d^2x \frac{\sqrt{-g_b}}{\kappa} \left[ \frac{1}{2} (Z_{,\mu} Z^{,\mu} - Y_{,\mu} Y^{,\mu}) - 2\lambda e^{Z/\xi} - e^{-Y/\xi} \frac{\Lambda_G}{2\lambda} F_{\mu\nu} F^{\mu\nu} + \xi (Z - Y) R_b \right]$$

with  $(g_b)_{\mu\nu} = g_{\mu\nu}|_{\varphi=0}$

This is a gauge independent equivalence!

Covers Liouville Gravity, the Witten Black Hole, the Callan-Giddings-Harvey-Strominger model and more

## Equations of motion

$$\text{Conformal gauge} \quad \left| \quad \text{Coulomb gauge} \right.$$

$$\lambda_0 = \lambda_1 = 1 \quad \left| \quad A_{,\mu}^\mu = 0 \rightarrow A_0 = A_{1,s} = 0 \right.$$

\* Liouville fields

$$Y_{,tt} - Y_{,ss} + \frac{\Lambda_G (A_{1,t})^2}{\lambda \xi} e^{-\frac{Y}{\xi}} = 0$$

$$Z_{,tt} - Z_{,ss} - \frac{2e^{\frac{Z}{\xi}} \lambda}{\xi} = 0$$

\* Gauge field current equations

$$\partial_\mu \left( A_{1,t} e^{-Y/\xi} \right) = 0 \rightarrow E = A_{1,t} e^{-Y/\xi} = \text{const}$$

\* Two constraints

$$(Y_{,t} \pm Y_{,s})^2 \mp 4\xi (Y_{,t} \pm Y_{,s})_{,s} + \frac{2\Lambda_G (A_{1,t})^2}{\lambda} e^{-\frac{Y}{\xi}} -$$

$$- (Z_{,t} \pm Z_{,s})^2 \pm 4\xi (Z_{,t} \pm Z_{,s})_{,s} + 4e^{\frac{Z}{\xi}} \lambda = 0 .$$

## Additional terms

- \* The Gibbons-Hawking-York boundary term has a LFT equivalent

$$\int_{\partial\mathcal{M}} dx \sqrt{\gamma} X K \quad \Longleftrightarrow \quad \int_{\partial\mathcal{M}} ds \, 2(Z - Y) (\xi K_b + \xi^2 (1 + 8g_b) Z_{,t})$$

- \* Minimally coupled *massless* scalar fields are trivially included

$$-\frac{1}{2} \sqrt{-g} \phi_{,\mu} \phi^{,\mu} \quad \Longleftrightarrow \quad -\frac{1}{2} \sqrt{-g_b} \phi_{,;\mu} \phi^{;\mu}$$



# Hamiltonian formulation

## Hamiltonian density & constraints

$$\mathcal{H} = \lambda_0 L^+ + \lambda_1 L^- + A_0 L^\emptyset$$

$$L^\pm = -\frac{1}{4} (P_Z \mp Z_{,s})^2 \mp \xi (P_Z \mp Z_{,s})_{,s} + \lambda e^{Z/\xi} + \\ + \frac{1}{4} (P_Y \pm Y_{,s})^2 \mp \xi (P_Y \pm Y_{,s})_{,s} + \frac{\lambda}{8\Lambda_G} e^{Y/\xi} \Pi_1^2 + \sum_{i=1}^D \frac{1}{4} (P_i \pm \phi_{i,s})^2$$

$$L^\emptyset = \Pi_{1,s}$$

## Classical algebra

$$\{L^\pm(f), L^\pm(g)\} = \pm L^\pm(fg' - f'g) \approx 0$$

$$\{L^+(f), L^-(g)\} = -\frac{1}{4\Lambda_G} \left( e^{Y/\xi} \Pi_1 L^\emptyset \right) (fg) \approx 0$$

## BRST formalism &amp; gauge fixing

- one pair of BRST ghosts per constraint
- BRST charge  $\{Q, Q\} = 0$
- BRST extensions:  $\mathcal{H}_{brst} = -\{\psi, Q\}$
- gauge fixing  $\lambda_0 = \lambda_1 = 1, A_0 = \alpha$

gauge fixed BRST extended Hamiltonian (density)

$$\mathcal{H}^{BRST} = L^{+,BRST} + L^{-,BRST} + \alpha L^{\emptyset, BRST}$$

Compactified space dimension  $s \in [0, 2\pi)$

e.o.m. admit chiral decomposition for the  $L^\pm$

Virasoro algebra

$$\{L_n^{\pm, BRST}, L_m^{\pm, BRST}\} = -i(n - m)L_{n+m}^{\pm, BRST}$$

## Quantization generalities

General á la Dirac quantization prescriptions:

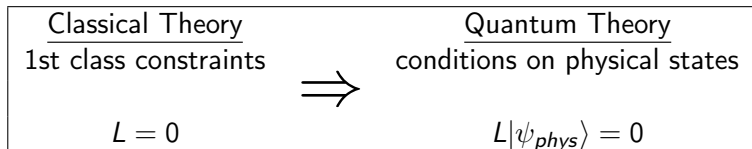
- Polarization of phase space
- Hilbert space
- Functions  $\rightarrow$  (normal ordered) Operators
- Poisson brackets  $\rightarrow$  Commutators  $\rightarrow$  central charges?
- 1st class (classical) constraints  $\rightarrow$  Conditions on physical states

Keeping in mind decoupling:

$$L^{\pm, BRST} = L^{\pm, Z} + L^{\pm, Y} + \sum_{i=1}^D L^{\pm, i} + L^{\pm, g}$$

we can quantize each sector separately

# Quantum realization of 1st class constraints



Determining  $\lambda$  for physical states

If  $|\psi\rangle$  is physical, what is the value of  $\lambda$ ?



$$\langle\psi_{phys}|L(\lambda)|\psi_{phys}\rangle = 0$$

Equations to be solved for  $\lambda$  and the parameters in  $\psi_{phys}$

# The quantum constraints

$$\begin{aligned}
 \langle L^+ + L^- \rangle &= -\frac{1}{2} \langle P_Z^2 + Z_{,s}^2 \rangle + \frac{1}{2} \langle P_Y^2 + Y_{,s}^2 \rangle + \sum \frac{1}{2} \langle P_i^2 + \phi_{i,s}^2 \rangle + \\
 &\quad + 2\xi \langle Z_{,ss} \rangle - 2\xi \langle Y_{,ss} \rangle + \\
 &\quad + \lambda \left[ 2 \langle e^{Z/\xi} \rangle + \frac{1}{4\Lambda_G} \langle e^{Y/\xi} \Pi_1^2 \rangle \right] + O(\hbar)
 \end{aligned}$$

$$\begin{aligned}
 \langle L^+ - L^- \rangle &= \langle P_Z Z_{,s} \rangle - 2\xi \langle P_{Z,s} \rangle + \langle P_Y Y_{,s} \rangle - 2\xi \langle P_{Y,s} \rangle + \\
 &\quad + \sum \langle P_i \phi_{i,s} \rangle + O(\hbar)
 \end{aligned}$$

$$\langle L^\emptyset \rangle = \langle \Pi_{1,s} \rangle + O(\hbar)$$

Compensating contributions from the  $Z$  and the  $Y$  sectors

# Ghost sector quantization

Chiral decomposition

$$c^\pm(s) = \sum_{n \in \mathbb{Z}} c_n^\pm e^{\mp ins} \quad b_\pm(s) = \sum_{n \in \mathbb{Z}} b_n^\pm e^{\mp ins}$$

with  $[c_n^\pm, b_m^\pm]^+ = \delta_{n+m}$

Radial quantization

$$\begin{aligned} [\hat{L}_n^{+,g}, \hat{L}_m^{+,g}] &= (n-m)\hat{L}_{m+n}^{+,g} - \frac{13}{6}(n^3-n)\delta_{m+n} \\ [\hat{L}_n^{-,g}, \hat{L}_m^{-,g}] &= (n-m)\hat{L}_{m+n}^{-,g} - \frac{13}{6}(n^3-n)\delta_{m+n} \\ [\hat{L}_n^{+,g}, \hat{L}_m^{-,g}] &= 0 \end{aligned}$$

# Liouville sector quantization

Quantum corrections to the coupling constants

$$\xi \rightarrow \xi_Z = \xi + \delta_Z \quad \xi \rightarrow \xi_Y = \xi + \delta_Y$$

$$L^\pm = \left[ -\frac{1}{4} (P_Z \mp Z_{,s})^2 \mp \xi_Z (P_Z \mp Z_{,s})_{,s} + \lambda e^{Z/\xi} \right] + \\ + \left[ \frac{1}{4} (P_Y \pm Y_{,s})^2 \mp \xi_Y (P_Y \pm Y_{,s})_{,s} + \frac{\lambda}{8\Lambda_G} e^{Y/\xi} \Pi_1^2 \right],$$

Mode expansions

$$Z(s) = \frac{i}{2\sqrt{\pi}} \left[ a_0 - a_0^\dagger + \sum_n' \frac{1}{n} (a_n e^{-ins} + \bar{a}_n e^{ins}) e^{-\epsilon|n|} \right], \\ P_Z(s) = \frac{1}{2\sqrt{\pi}} \left[ a_0 + a_0^\dagger + \sum_n' (a_n e^{-ins} + \bar{a}_n e^{ins}) e^{-\epsilon|n|} \right],$$

Commutation relations

$$[a_n, a_m] = [\bar{a}_n, \bar{a}_m] = n\hbar \delta_{-m}^n, \quad a_n^\dagger = a_{-n}, \quad \bar{a}_n^\dagger = \bar{a}_{-n}, \quad [a_0, a_0^\dagger] = \hbar.$$

Quantum Virasoro algebra in the Liouville sectors

$$[L_r^{\pm, Z}, L_q^{\pm, Z}] = (r-q)\hbar L_{r+q}^{\pm, Z} + \hbar \left[ \left( \frac{\hbar}{24} - 4\pi \left( \xi - \frac{\hbar}{8\pi\xi} \right)^2 \right) r^3 + \hbar \frac{1}{12} r \right] \delta_{r+q}$$

$$[L_r^{\pm, Y}, L_q^{\pm, Y}] = (r-q)\hbar L_{r+q}^{\pm, Y} - \hbar \left[ \left( \frac{\hbar}{24} - 4\pi \left( \xi - \frac{\hbar}{8\pi\xi} \right)^2 \right) r^3 + \hbar \frac{1}{12} r \right] \delta_{r+q}$$

with the choices

$$\xi_Z = \xi_Y = \xi - \frac{\hbar}{8\pi\xi}$$



## Additional fields

The Y field with no Maxwell-field

$$L^{\pm, \Upsilon} = \left[ \frac{1}{4} (P_{\Upsilon} \pm \Upsilon_{,s})^2 \mp \xi_{\Upsilon} (P_{\Upsilon} \pm \Upsilon_{,s})_{,s} \right]$$

No need to fix  $\xi_{\Upsilon}$

$$\left[ L_r^{\pm, \Upsilon}, L_q^{\pm, \Upsilon} \right] = (r - q) \hbar L_{r+q}^{\pm, \Upsilon} - \hbar \left[ \left( \frac{\hbar}{24} - 4\pi \xi_{\Upsilon}^2 \right) r^3 + \frac{\hbar}{12} r \right] \delta_{r+q}$$

Massless scalar fields

$$\left[ L_r^{\pm, \phi}, L_q^{\pm, \phi} \right] = (r - q) \hbar L_{r+q}^{\pm, \phi} - \hbar^2 \left( \frac{1}{24} r^3 + \frac{1}{12} r \right) \delta_{r+q}$$

## Quantum Virasoro algebra: the case with no Maxwell field

$$L^{\pm, BRST} = L^{\pm, Z} + L^{\pm, \Psi} + L^{\pm, g} + \sum_{i=1}^D L^{\pm, \phi_i}$$

Central charge

$$c(\xi_{\Psi}) = \hbar \left[ \left( -4\pi \left( \xi - \frac{\hbar}{8\pi\xi} \right)^2 + \hbar \frac{52 - D}{24} + 4\pi\xi^2 \right) r^3 + \hbar \frac{26 - D}{12} r \right]$$

Eliminated by fixing  $\xi_{\Psi}$  (with a possible constraint on  $D$  depending on the value of  $\xi$ ) and shifting the zero modes

$$L_0^{\pm, BRST} \Rightarrow L_0^{\pm, BRST} - \hbar^2(D - 26)/24$$

## Which basis/representation?

- Fock excitations for the states (dropping ghosts contributions)  
*allows choices of the states to "test"*

$$|\psi^f(d)\rangle = \bigotimes_{n \in \mathbb{Z}} \left[ \sum_{\mu \geq 0} d_{\mu}^f(n) |\mu_n^f\rangle \right]$$

$$|\psi\rangle = \sum_{\{d^Z, d^Y, d^i\}} |\psi^Z(d^Z)\rangle |\psi^Y(d^Y)\rangle \bigotimes_{i=1}^D |\psi^i(d^i)\rangle$$

- Coherent states basis for operators  
*diagonal integral representation of operators*

$$\hat{O} = \int \prod_m \left[ \frac{d^2 z_m}{2\pi} \right] |\underline{z}\rangle \varphi(z, \bar{z}) \langle \underline{z}| \quad |\underline{z}\rangle = \bigotimes_f \left( \bigotimes_n |z_n^f\rangle \right)$$

$$\varphi(z, \bar{z}) = e^{(\sum_m \partial_{z_m} \partial_{\bar{z}_m})} \langle \underline{z}| \hat{O} | \underline{z}\rangle$$

- Projection on coherent states:  $\langle n|z\rangle = \langle \Omega|a^n|z\rangle = z^n e^{-\frac{1}{2}|z|^2}$
- Subset of states with factorized wave functions:  
 $|\psi|^2 = |\psi^Z|^2 |\psi^Y|^2 |\psi^1|^2 \dots$
- The constraints are a sum of gaussian integrals over complex variables:

$$\begin{aligned}
 \langle L^\pm(s) \rangle &= \int \prod_m \left[ \frac{dz_m d\bar{z}_m}{2\pi} \right] t_Z^\pm(s, z, \bar{z}) |\psi^Z(\underline{z})|^2 + \\
 &+ \int \prod_m \left[ \frac{dz_m d\bar{z}_m}{2\pi} \right] t_Y^\pm(s, z, \bar{z}) |\psi^Y(\underline{z})|^2 + \\
 &+ \sum_i^D \int \prod_m \left[ \frac{dz_m d\bar{z}_m}{2\pi} \right] t_i^\pm(s, z, \bar{z}) |\psi^i(\underline{z})|^2
 \end{aligned}$$

## The vacuum

The classical cosmological constant needs to vanish

At the quantum level

$$\langle L^+ + L^- \rangle = -\lambda + \frac{\hbar^2}{24} (26 - D) \quad \langle L^+ - L^- \rangle = 0$$



$$\lambda = \frac{26 - D}{24} \hbar^2$$

Quantum correction to  
the classical cosmological constant  
given by the (reabsorbed) central charge of the quantum algebra  
it 'knows' the matter content of the theory

# First level excitations

Quantum states

$$|1\rangle = \bigotimes_f^{\text{fields}} |1^f\rangle = \bigotimes_f^{\text{fields}} \left[ \bigotimes_{n \neq n^f} | \Omega \rangle \right] \otimes \left[ d_0^f(n^f) | \Omega \rangle + d_1^f(n^f) | 1_{n^f}^f \rangle \right]$$

Non-zero modes constraints

$$\begin{aligned} \langle L_n^+ + L_{-n}^- \rangle &\propto 2\lambda \left( \bar{d}_0^Z(n) d_1^Z(n) + d_0^Z(-n) \bar{d}_1^Z(-n) \right), \\ \langle L_n^+ - L_{-n}^- \rangle &= \xi_{\mathcal{Y}} \left( \bar{d}_0^{\mathcal{Y}}(n) d_1^{\mathcal{Y}}(n) - d_0^{\mathcal{Y}}(-n) \bar{d}_1^{\mathcal{Y}}(-n) \right) - \\ &\quad - \xi_Z \left( \bar{d}_0^Z(n) d_1^Z(n) - d_0^Z(-n) \bar{d}_1^Z(-n) \right), \end{aligned}$$

Only pure excitations for  $Z$  and  $\mathcal{Y}$

$$d_0^Z(n^Z) = d_0^{\mathcal{Y}}(n^{\mathcal{Y}}) = 0$$

## Zero modes constraints

$$\begin{aligned} \langle L_0^+ + L_0^- \rangle &= \hbar^2 \frac{26 - D}{12} - 2\lambda \left( 1 + \frac{|d_1^Z(n^Z)|^2}{4\pi\xi^2|n^Z|} \Big|_{n^Z \neq 0} \right) \\ &\quad + \frac{1}{4\pi} \sum_f^{\text{fields}} \left[ \beta(f) \left( 2|n^f| + \delta_0^{n^f} \right) |d_1^f(n^f)|^2 \right] \\ \langle L_0^+ - L_0^- \rangle &\propto \sum_f^{\text{fields}} \left[ \beta(f) n^f |d_1^f(n^f)|^2 \right] \end{aligned}$$

with  $\beta = 1$  for matter fields and  $\Psi$ , and  $\beta = -1$  for  $Z$

# D=0 first level excitations

- ① Both fields are excited,  $n^Z = n^Y = N$

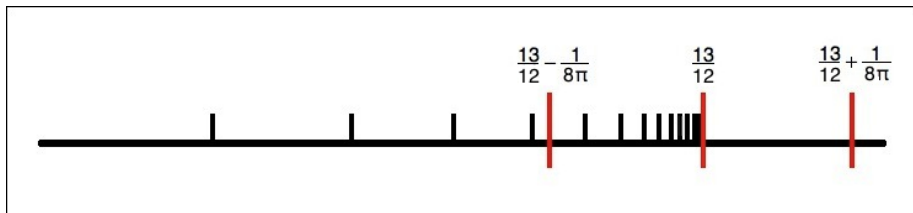
$$\lambda = \hbar^2 \frac{13}{12} \left( 1 + \frac{1}{4\pi\xi^2 |n^Z|} \Big|_{n^Z \neq 0} \right)^{-1} = \begin{cases} \hbar^2 \frac{13}{6} \frac{2\pi\xi^2 |N|}{1+4\pi\xi^2 |N|} & : N \neq 0, \\ \hbar^2 \frac{13}{12} & : N = 0. \end{cases}$$

- ② Only the Z field is excited  $n^Z = 0$

$$\lambda = \hbar^2 \frac{13}{12} - \frac{1}{8\pi}$$

- ③ Only the  $B^0$  field is excited  $n^Y = 0$

$$\lambda = \hbar^2 \frac{13}{12} + \frac{1}{8\pi}$$





## Higher excitations

$D > 0$  gives loosened constraints  $\Rightarrow d$ 's not fixed

- \* 1st level excitations for  $D = 1$ 
  - the spectrum is bounded from below (the minimum depends on  $\xi$ )
  - unbounded towards  $+\infty$
  - discrete values + continuous band
- \* Higher excitations:
  - main difficulty: the  $s$ -dependence cannot be eliminated with a Fourier transform
  - main interest: contributions from  $\xi_Z$ ,  $\xi_{\mathcal{Y}}$  to  $\lambda$
  - high excitations of  $Z$  compensate arbitrary excitations for  $\mathcal{Y}$  and scalar fields

## Summary and conclusions

- in 1+1 dimensions a gauge independent duality exists between a class of models of dilaton-Maxwell gravity and Liouville Field Theory
- the cosmological constant is fixed by space-time symmetry in the quantum theory
- the quantum gravitational d.o.f. contributes with a negative term VS positive matter contributions
- the cosmological constant spectrum is 'built' around the central charge of the Virasoro algebra

Thank you!

References [arXiv:1102.4957](https://arxiv.org/abs/1102.4957) [arXiv:1111.1612](https://arxiv.org/abs/1111.1612)